

Heat transport measurements in turbulent rotating Rayleigh-Bénard convection

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We present experimental heat transport measurements of turbulent Rayleigh-Bénard convection with rotation about a vertical axis. The fluid, water with a Prandtl number (σ) of about 6, was confined in a cell with a square cross section of 7.3×7.3 cm² and a height of 9.4 cm. Heat transport was measured for Rayleigh numbers $2 \times 10^5 < Ra < 5 \times 10^8$ and Taylor numbers $0 < Ta < 5 \times 10^9$. We show the variation in normalized heat transport, the Nusselt number, at fixed dimensional rotation rate Ω_D , at fixed Ra varying Ta, at fixed Ta varying Ra, and at fixed Rossby number Ro. The scaling of heat transport in the range of 10^7 to about 10^9 is roughly 0.29 with a Ro-dependent coefficient or equivalently is also well fit by a combination of power laws of the form $a Ra^{1/5} + b Ra^{1/3}$. The range of Ra is not sufficient to differentiate single power law or combined power-law scaling. The data are roughly consistent with an assumption that the enhancement of heat transport owing to rotation is proportional to the number of vortical structures penetrating the boundary layer. We also compare indirect measures of thermal and Ekman boundary layer thicknesses to assess their potential role in controlling heat transport in different regimes of Ra and Ta.

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I. INTRODUCTION

Turbulent thermal convection plays a key role in many of the phenomena associated with geophysical and astrophysical fluid dynamics [1] as well as providing a well-posed problem for the study of fundamental fluid dynamics [2]. In several important examples including oceanic deep convection [1] and convection in stars [3] and giant planets [4], the effects of rotation are critical in determining the nature of the fluid motion. Rotation also provides an additional parameter for understanding the origins of heat transport scaling in turbulent convection, a topic of tremendous experimental activity in recent years [2,5]. In comparison, the research efforts applied to rotating turbulent convection have been rather modest arising from the pioneering theoretical work of Chandrasekhar [6,7]. Experimental measurements of heat transport in rotating convection include the seminal work of Rossby [8] and later studies that also had qualitative flow visualization [9]. Numerical simulations have also had significant impact [10–14]. Here, we consider both rotating and nonrotating convection and provide insights into heat transport scaling of rotating convective turbulence. A short report of some aspects of this work appeared previously [15], and further studies of velocity fields in rotating convection motivated by this work were also published [16]. Recent experimental work has elucidated the dependence of heat transport on Prandtl number as well as on Rayleigh and Rossby numbers [17], made measurements of velocity and vorticity fields [18], and related the crossover to rotation-dominated convection through an analysis in terms of Ekman boundary layers [19].

Rotating Rayleigh-Bénard convection can be characterized by three dimensionless parameters: the Rayleigh num-

ber Ra which is a measure of buoyant forcing, the Taylor number Ta which measures the effect of the rotational Coriolis force, and the Prandtl number σ which determines the dominant nonlinearity in convection. These parameters are defined by

$$Ra = \frac{g\alpha d^3 \Delta T}{\nu\kappa}, \quad Ta = \left(\frac{2\Omega_D d^2}{\nu} \right)^2, \quad \sigma = \frac{\nu}{\kappa}, \quad (1)$$

where g is the acceleration of gravity, α is the thermal expansion coefficient, ΔT is the temperature difference across the fluid layer of height d , ν is the kinematic viscosity, κ is the thermal diffusivity, and Ω_D is the physical angular rotation frequency. Properties of thermal turbulence can also be affected by the cell geometry characterized by the ratio of a lateral length to a vertical length. For our square geometry, we define the cell aspect ratio as $\Gamma \equiv L/d$, where L is the lateral size of the cell.

Although Ra, Ta, σ , and Γ completely define the parameter space of rotating convection, the behavior of different quantities such as heat transport is complicated when one parameter is varied while keeping the others constant. For example, as Ra is increased at fixed Ta, the relative influence of buoyancy and rotation changes, making an evaluation of the influence of rotation alone difficult. To mitigate this effect, it is useful to define [13] a different measure of rotation—the convective Rossby number Ro

$$Ro = \sqrt{\frac{Ra}{\sigma Ta}} \quad (2)$$

which is a ratio of a buoyancy time scale $\tau_b \sim \sqrt{d/(g\alpha\Delta T)}$ to a rotational time $\tau_r \sim 1/\Omega_D$. This definition is equivalent to those used previously [11,20] and is closely related to other definitions of convective Rossby number [10,21,22]. Roughly speaking, the border between rotation-dominated and buoyancy-dominated flows should be approximated by the condition $Ro = 1$.

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The quantity of interest here is turbulent heat transport as measured by the Nusselt number Nu , which is the total heat transported by convection normalized by the heat transported by molecular diffusion alone. To appreciate the influence of rotation on Nu , it is important to understand the dependence of Nu on Ra without rotation. The investigation of nonrotating convection has been extensive over almost 40 years with early work focused on classical theories [23–25] that predicted a power-law relationship of the form $Nu = A Ra^\beta$ with $\beta = 1/3$. Later measurements, particularly those in helium gas, suggested a value $\beta = 2/7$ with theory and early numerical simulations providing a solid basis for such a law. A detailed review of these results was presented by Siggia [2]. An extension [26,27] of the kinetic and thermal boundary layer theory [28] that included an expanded analysis of competing boundary and bulk dissipation processes produced a phase diagram with power-law exponents of simple integer ratios, i.e., $1/5$, $1/4$, $1/3$, and $1/2$, in each region. Because of the finite extent of each region, one expects crossover effects between different regions which suggests a form $Nu = a Ra^{\beta_1} + b Ra^{\beta_2}$ with specific predictions for the coefficients derived from fitting a few data sets in each region: for $\sigma \approx 6$, the predictions are $\beta_1 = 1/5$ and $\beta_2 = 1/3$ [27]. In this latter regard, high-precision experimental data for room-temperature fluids [29,30] have been extremely valuable in elucidating differences between a single power-law description and one involving two power laws with fixed coefficients. The measurements presented here are over a modest range of $Ra < 10^9$ and, thus, cannot distinguish between these two forms. We compare our results with other measurements of heat transport of nonrotating convection as a benchmark for considering the effects of rotation on turbulent heat transport.

The effects of rotation on convection, especially on heat transport, might be expected to be substantial given that rotation profoundly changes the nature of boundary layer instability and modifies the length scales over which motions occur. Whereas thermal plumes are formed in long sheets and are swept across the cell by mean flow, rotation spins up these plumes into intense vortical structures. Furthermore, rotation is known to shorten the linear length scale dramatically as rotation is increased [7]. Additional ingredients introduced by rotation are the Ekman pumping or suction imposed by the differential rotation of the boundary and the interior flow and the dynamical constraints imposed by the Taylor-Proudman theorem for strongly rotating flows. Despite these interesting factors, previous heat transport measurements have not been well understood for a number of reasons. Rossby, in his seminal paper on rotating convection [8], reported heat transport measurements for water and for mercury as a function of Ra and Ta with emphasis on the regions close to onset and of moderate Ra ($< 3 \times 10^6$). His measurements, as well as later measurements in helium [31], quantitatively showed that the convective onset was below the theoretical prediction of linear stability analysis [6,7]. This reduction in critical Rayleigh number was attributed to a transition to azimuthally periodic modes localized near the wall [31,32] but neither of the experiments had flow visualization capabilities. The existence of such wall states was later confirmed [9,33] using shadowgraph flow visualization,

but rather than being stationary the sidewall states were observed to precess in the rotating frame counter to the direction of rotation. This resolved one difficulty with the data set of Rossby. In what follows, we denote Ra_c (Ra_b) as the onset Ra for the wall-mode state (bulk state).

Other experiments [21,34,35] and numerical simulations [10,12,22] of rotating convection involve an open upper fluid surface where one can visualize the development of convective structures and the interaction of vortices and characterize some of the statistics of the temperature and velocity fields. These experiments are, however, not amenable to the measurement of accurate heat transport.

The heat transport experiments [9] that motivated this work used water in a cylindrical cell [36] with top and bottom rigid boundaries, and measurements were made to higher Ra ($\approx 2 \times 10^7$) than Rossby [8]. The normalized heat transport Nu at a constant rotation rate appeared, however, to asymptote to the nonrotating result at high Ra and did not exhibit a clear power-law scaling over the limited range of Ra . As noted above, maintaining a fixed Rossby number mitigates against the crossover to buoyancy dominated flow at high Ra and allowed numerical simulations [14] to show that Nu scaled approximately as the $2/7$ power for a fixed convective Rossby number with no-slip top and bottom boundary conditions. To test this prediction and to further characterize heat transport as a function of rotation, it was necessary to extend the heat transport measurements to higher Ra than in [9].

This paper is organized as follows. The experimental apparatus and procedures are described in Sec. II. The heat transport results of nonrotating and rotating turbulence are presented in Secs. III and IV, respectively. Sec. V summarizes the paper.

II. EXPERIMENTAL PROCEDURE

A. Rotating apparatus and cell

The experimental apparatus, shown schematically in Fig. 1, was an improved version of the one used previously in studies of rotating convection [9,37]. The top-plate temperature of the cell was controlled by water flow which was distributed evenly by a set of 12 turrets divided into two groups pointing to $1/3$ and $2/3$ of the radius, respectively.

The water flow was temperature controlled by a refrigerator circulator and then fed into the rotating frame through a water slip connection. The flow was further temperature regulated by a feedback temperature control unit that maintained the top-plate temperature constant with root-mean-square fluctuations of less than 1 mK. A film heater attached to the bottom plate provided a constant heat current to the fluid layer. The power input was obtained by measuring the voltage across the film heater and the current through it. The latter was obtained from the voltage across a temperature-controlled standard resistor.

All electrical wires were fed into the rotating frame through an electrical slip ring and inside a hollowed steel shaft which also served as the drive train for rotation. The electrical noise of the slip ring was small enough that there was no measurable difference in the signals with or without

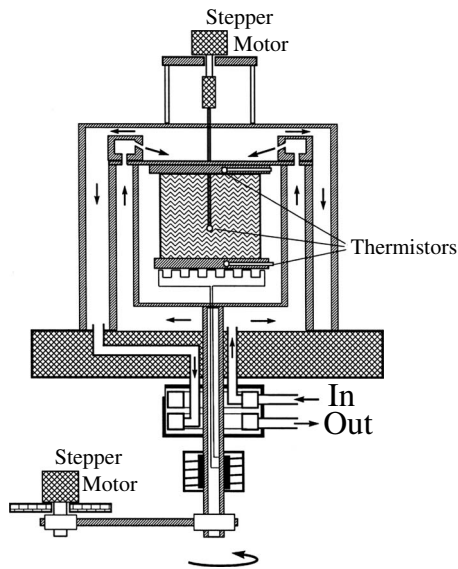


FIG. 1. Schematic of the experimental apparatus.

rotation. Rotation was provided by a microstepper motor through a shaft, two gears, and a timing belt and was under computer control. The gear ratio sets a lower limit for the frequency of about 0.01 Hz. The maximum frequency surveyed was 0.5 Hz.

The convection cell was constructed with aluminum top (1.27 cm thick) and bottom (0.64 cm thick) plates and Plexiglas sidewalls (0.32 cm thick). The aluminum plates were anodized to prevent corrosion in water. The cell had a height $d=9.40$ cm and horizontal dimensions $L_x=L_y=7.30$ cm with an aspect ratio of $\Gamma=L_x/d=0.78$. The cell's bottom and sides were insulated by 2.5-cm-thick Styrofoam to reduce thermal losses from radiation and conduction or convection by air. Four thermistors were embedded in the top plate at positions $\{x,y\}=\{L_x/4,L_y/2\},\{3L_x/4,L_y/2\},\{L_x/2,L_y/4\},\{L_x/2,3L_y/4\}$ with the vertical centerline of the thermistor at 0.32 cm from the fluid. The four bottom thermistors were positioned similarly but at a vertical distance of 0.23 cm from the fluid. The four thermistors in each plate gave the average plate temperature. The wires from the film heater and from the four thermistors were heat sunk on the bottom of the cylindrical can, which was maintained at the same temperature as the top plate.

At the top of the apparatus, a thermistor probe could be inserted into the cell to measure the temperature at a point (or small number of points) within the cell. The probe was a 0.32 cm stainless steel tube that penetrated the top plate through a snug fit hole at the plate center. The height of the probe was adjusted by a stepper motor that was under computer control. Using this probe one could measure the mean vertical temperature profile including the thermal boundary layer structure. One probe had a single thermistor along the cell centerline whereas a second probe had five thermistors positioned along a line spanning about 70% of the diagonal distance in the cell to allow measurement of lateral variations in the temperature distribution. Details of the temperature measurement and results on the local measurements of the temperature field and thermal boundary layer will be presented elsewhere.

B. Heat transport

Heat is transported more efficiently by convection, where heat can be advected by the fluid motion, than by conduction where heat is transported solely by molecular diffusion. The enhancement of thermal transport by convection is characterized by the Nusselt number $Nu=K_{eff}/K$, where K_{eff} and K are the effective thermal conductance and molecular thermal conductance of the fluid layer, respectively. In an experimental realization of convection, heat \dot{Q} is applied to the bottom plate while the top plate is maintained at constant temperature T_t . Most of the applied heat flows through the fluid but a small amount is lost to parasitic thermal conductance including heat flow through the cell side walls, thermal conduction along other pathways such as wires or foam insulation surrounding the cell, and thermal radiation between the cell bottom plate or sidewalls and surfaces at different temperatures. An accurate measurement of fluid heat transport requires accounting for these different parasitic conduction pathways and properly subtracting them from the total measured heat transport. Below the onset of convection, the background conductivity can be measured and effectively subtracted from the total heat transport contribution provided that the thermal conductivity of water, available from the literature, is assumed. For turbulent convection in room-temperature experiments, this is difficult to accomplish because if $\Delta T \approx 10$ K corresponds to the maximum achievable $Ra \approx 10^9$, then the onset of convection occurs at $\Delta T_c \approx 2 \times 10^{-5}$ K, far below experimental resolution. This problem can be overcome by rotating the cell, thereby increasing the onset of convection by about four orders of magnitude [7] (note $\Delta T_c \sim Ra_c \sim Ta^{2/3}$), so that $\delta T_c \approx 0.2$ K, thereby allowing the background thermal conductivity to be measured directly. We now describe this procedure in detail.

The background thermal conductance was determined to be $K_b=0.0423 \pm 0.0004$ W/K, comparable to the water layer's thermal conductance $K=0.0341$ W/K at a cell mean temperature of 21.5 °C. Assuming that K_b is independent of the mean temperature of the cell, the Nusselt number is given by

$$Nu = \frac{\dot{Q}/\Delta T - K_b}{K}, \quad (3)$$

where $\Delta T=T_b-T_t$ is the temperature difference across the water layer with T_t and T_b being the top-plate and the bottom-plate temperatures and K is the thermal conductance of the water layer at the mean temperature $T_0=(T_t+T_b)/2$. Subtracting the sidewall conductance is important in these measurements because it can account for as high as a 10% correction at the smallest Nu measured for the purposes of determining scaling behavior, which is $Nu > 15$. The top and the bottom temperatures are corrected for the temperature drop in the aluminum plates although this correction was always less than 0.3%.

In our experiment, we measured K_b at a single temperature because the temperature dependence of the background terms is quite small and does not affect the data presented here. To evaluate the systematic error in our measurements of K_b , we have estimated the different contributions in that

quantity. The major contributors to K_b are the Plexiglas sidewalls, the insulating foam, the electrical wires, and thermal radiation. The first three contributions have very weak temperature dependence (less than 1% change over a 20 K temperature difference) and small magnitude, estimated to be 0.006, 0.006, and 0.003 W/K, respectively. The rest of the measured background heat transport, about 60% of K_b (≈ 0.027 W/K), comes from a series combination of conduction through the foam, convection in the air surrounding the insulation, and from thermal radiation from the outer surfaces of the insulation to the surroundings, which are maintained at the top-plate temperature. The top-plate temperature was held constant and the bottom-plate temperature was changed by a maximum of 20 K corresponding to the maximum heat input. This produces about a 2 K increase in the average temperature of the radiating surfaces, or about a 2% increase in radiated heat, which results in about a 0.02 overestimate in Nu. Therefore, for the measurements with fixed T_t , K_b can be taken to be constant as the correction to Nu is less than 0.1%.

This analysis neglects an important point regarding the heat transported through the sidewalls [38]. Rather than supporting a linear temperature profile as in the nonconvecting state where the background is measured, the sidewalls are in contact with a turbulent fluid that is approximately isothermal in the bulk of the flow. For nonrotating convection this effect leads to an enhancement of heat transport through the sidewalls in the turbulent state. For conditions similar to those presented here (thin Plexiglas walls relative to the lateral extent of the system), however, the correction to the total heat transport was shown by numerical modeling [38] to be small, ranging from about 2% for $Nu=10$ to 1% at $Nu=100$. Our sidewall is thicker than the one used in the numerical model by about a factor of 2, so in the worst case these values might be twice as large. Applying a correction of this order shifts the exponent of a power-law fit by at most 1% (higher) and the constant term by about 7% (lower). These estimates contribute to the systematic error in our results but we do not explicitly correct for the sidewall effect in the data presented below. To properly account for the sidewall in a systematic way for rotating convection, one would need to take into account the mean vertical temperature gradient that develops in rotating thermal convection [13,39].

C. Parameter space

The parameter space for rotating convection is defined by Ra and Ta, which are proportional to the physical control variables of ΔT and Ω_D , respectively. In Fig. 2, the parameter space is shown over a range which encompasses our experimental measurements. The shaded region denotes the area in which most of our efforts are focused.

The limitations which determine that area are the range of measurable ΔT for a given cell height d ; a four-decade variation in Ra was obtained by varying the temperature difference across the cell from 2 mK to 20 K. The Taylor number for the shaded area ranged from about 1×10^6 to 5×10^9 , corresponding to rotation frequencies from 0.01 to 0.5 Hz. The lower limit was determined by the range of stability of

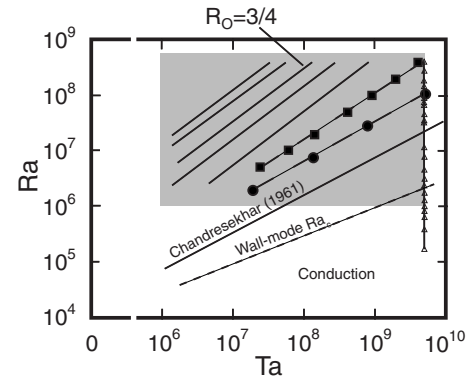


FIG. 2. Parameter space of Ra vs Ta. Most measurements were conducted in the gray area by fixing Ra, Ω_D , or Ro. The measurements at $\Omega_D=0.0$ and 3.14 rad/s ($Ta \approx 5.0 \times 10^9$) started at Ra as low as 5×10^5 . Five solid lines in the gray area represent five different Ro (from right to left, 0.30, 0.52, 0.75, 1.15, and 1.49) used in the measurements. Data (●), from heat transport measurements at constant Ω_D , set the lower bound of Ra above which Nu could be expected to exhibit turbulent convection. Symbols (■), deduced from Nu data at fixed Ra, represent the loci of maximum Nu at constant Ra. The data (▲) at the highest $Ta \approx 5 \times 10^9$ spanned the largest Ra number range. The theoretical predictions of bulk [7] and wall-mode [40,41] convective onsets under rotation are shown as solid and dashed lines, respectively, and labeled in the figure.

the stepping motor given a particular gear ratio. By reducing that ratio, lower Taylor numbers could have been investigated but this proved unnecessary as the interesting range of Ta was spanned with the chosen gear ratio.

During the experiments, we fixed some parameters and studied the dependence of measured quantities on the others. For instance, we fixed Ra to study the dependence on Ta, and vice versa. One important parameter, the convective Rossby number Ro which provides a relative measure of buoyancy relative to rotation, was maintained fixed by varying both ΔT and Ω_D for each data point. The contours of constant Ro, plotted in Fig. 2, are approximately straight lines in the log-log plot (small deviations arise from variations in σ with Ra of about 20% over the full Ra range, discussed in detail below). Numerical simulations of turbulent rotating convection with $\sigma=1$ [13,14] followed the contour of $Ro=3/4$ where buoyancy and rotation had roughly comparable importance. This particular line is noted in the figure. Several other sets of data shown in the figure are discussed later. Overall, the parameter ranges in our experiments fall roughly into the region studied in open-top experiments [21,34,35]; the parameter space of experiments on rotating convection prior to about 1990 was summarized in [21]. To compare between the two parameter spaces, one needs to notice that the flux Rayleigh number used in the open-top experimental parameter space is related to Ra by $Ra_f = Nu Ra$ so that the highest Ra in our experiments, where $Nu \approx 60$ corresponds to $Ra_f \approx 3 \times 10^{10}$.

III. HEAT TRANSPORT IN NONROTATING CONVECTION

In this section, we present experimental results for nonrotating convection. This enables us to compare our results

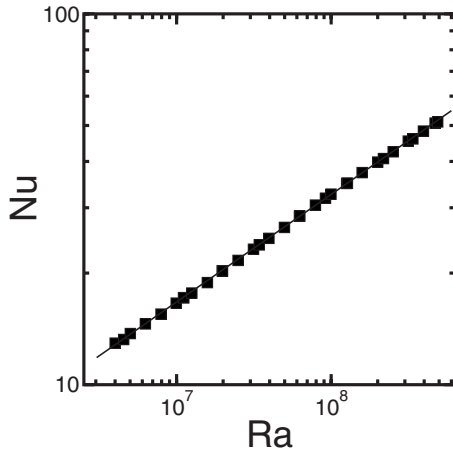


FIG. 3. Nu vs Ra for nonrotating convection. The solid line is the power-law fit $\text{Nu}=0.158 \text{ Ra}^{0.29}$ over the range $4 \times 10^6 < \text{Ra} < 5 \times 10^8$. An equivalent fit is $\text{Nu}=0.26 \text{ Ra}^{1/5} + 0.047 \text{ Ra}^{1/3}$.

with existing theories and with other nonrotating convection experiments, of which there are many. It also serves as a reference for our results on rotating convection. We concentrate here on measurements of heat transport in fluids with Prandtl number $\sigma \approx 6$.

The Nusselt number of nonrotating convection measured with fixed T_t is shown in Fig. 3 as a function of Ra for $4 \times 10^6 < \text{Ra} < 5 \times 10^8$. The results are reasonably well described by a power law with scaling coefficients $A = 0.158 \pm 0.003$ and $\beta = 0.289 \pm 0.002$, which were obtained by fitting all data for $4 \times 10^6 < \text{Ra} < 5 \times 10^8$ shown in the figure. Fitting the data between 4×10^7 and 5×10^8 yields slightly different values: $A = 0.164 \pm 0.003$ and $\beta = 0.286 \pm 0.002$. These latter values have the virtue of being derived from a range of Ra that is more fully in the turbulence regime but with the disadvantage of a shorter scaling range. To account for this systematic uncertainty we take $A = 0.164 \pm 0.006$ and $\beta = 0.286 \pm 0.003$ as the best estimates for the scaling coefficients. The exponent is very close to $2/7$ and agrees well with earlier work as summarized in Table I.

Although the $2/7$ value has been shown not to describe heat transport data over a much larger range of Ra [29], we use it here for convenience. Further, the uncertainties associated with the exponents and coefficients are the result of statistical fits and underestimate the systematic errors associated with the different experiments.

Compared to the scaling exponent β , the coefficient A is quite different from one experiment to another. A is sensitive to the exponent and a precise determination of A requires a larger range of Ra than has been available in any of the experiments using water. Fixing the exponent at $2/7$ and computing a value for $A_{2/7}$ (equivalent to $\text{Nu}/\text{Ra}^{2/7}$) at different Ra give a better comparison between data sets (see Table I). For the water experiments and for $10^7 < \text{Ra} < 10^9$, the coefficients agree quite well except for the experiments of Solomon and Gollub [44] where a liquid mercury bottom surface may account for the discrepancy. All of the data reported earlier and listed in Table I do not directly measure the background heat transport contribution that we are able to account for using rotation. This background measurement is important in eliminating systematic error for smaller Nu. An average over all the data sets for convection in water yields $A = 0.161 \pm 0.007$ and $\beta = 0.287 \pm 0.008$ with no statistically significant dependence of Nu on Γ . In summary, our data for nonrotating convection agree well with earlier results despite the significant variation in aspect ratios between experiments.

We also took heat transport data at fixed mean-cell temperature T_0 to estimate the Prandtl number dependence of the heat transport [15] (not reported here). We can use that data to correct the data at fixed T_t . In the fixed T_t measurements reported in the remainder of this paper, we had a variation in the range $21.5 < T_0 < 31.4$ °C (and resultant variation $6.7 > \sigma > 5.2$) corresponding to changes from the lowest to highest heat input. Such nonconstant σ or T_0 results in an uncertainty in Nu on the order of 0.8%. We interpolate the Nu data to a constant mean temperature or a constant σ . We choose $T_0 = 26.0$ °C (where $\sigma = 5.93$) as the reference temperature which was about the average of the mean temperatures in the experiments. The interpolated value is given

TABLE I. Values of heat transport scaling parameters: A , β , $A_{2/7}$ ($\text{Ra} = 10^7$), $A_{2/7}$ ($\text{Ra} = 10^8$), Ra range, Γ (C: cylindrical with $\Gamma = \text{diameter}/\text{height}$, S: square with $\Gamma = \text{width}/\text{height}$, R: rectangular with $\Gamma_x \times \Gamma_y$), and reference. All the experiments listed here used water as the working fluid. The Prandtl number for different experimental conditions varied slightly but was in the range $4 < \sigma < 7$.

A	β	$A_{2/7}$ (10^7)	$A_{2/7}$ (10^8)	Ra range ($\times 10^6$)	Γ	Reference
0.131	0.300(5)	0.165	0.170	0.03–2	2.8–10 (C)	Rossby (1969) [8]
0.183	0.278	0.162	0.159	0.3–100	1.5 \times 2.5(R)	Chu and Goldstein (1973) [42]
0.145	0.29	0.155	0.156	30–4000	3.5–14 (S)	Tanaka and Miyata (1980) [43]
0.137	0.275(7)	0.115	0.112	2–200	0.71, 1.6 (S)	Solomon and Gollub (1991) [44]
0.129	0.299(3)	0.160	0.164	0.1–20	2.0 (C)	Zhong <i>et al.</i> (1993) [9]
0.19(4)	0.28(1)	0.172	0.170	1–400	4 \times 1 (R)	Chillá <i>et al.</i> (1993) [45]
0.145	0.292	0.160	0.162	400–7000	1.0 (C)	Cioni <i>et al.</i> (1996) [46]
0.16	0.281(2)	0.147	0.145	800–6000	1.0 (C)	Shen <i>et al.</i> (1996) [47]
0.164(6)	0.286(3)	0.164	0.164	40–500	0.78 (S)	This work
0.154	0.291	0.168	0.170	3000–60000	1.0 (C)	Nikolaenko and Ahlers (2003) [30]

by $Nu = Nu_m(\sigma/5.93)^{0.029}$, where Nu_m is the measured Nusselt number using Eq. (3). The difference between the interpolated and the measured values of Nu is less than 0.8% and does not change the scaling coefficients within their specified error bars. Nevertheless, the interpolated values are reported in Fig. 3.

One important feature of turbulent convection is the large-scale circulation that is driven by an accumulation of thermal plumes that congregate near the lateral boundaries [2,5]. The general circulation in our cell was visualized using glass encapsulated thermochromic liquid crystals (TLCs). A white light sheet about 0.1 cm in width was used to illuminate the cell from the side and a black background was provided for good contrast. The mean-flow direction was typically across the cell diagonal. Flow reversals were observed as was a shifting of the main diagonal circulation in a clockwise or counterclockwise direction as viewed from above. Occasionally, the flow would shift to the other cell diagonal. This cross diagonal flow has been observed before in convection cells with square cross section [44]. In addition to the large diagonal flows there were often small recirculating cells in the corners and along the bottom-side boundary. Viewed from above the thermal plumes near the bottom boundary layer were arranged into coherent sheets which were swept up by the mean flow. This was also seen in a number of convection experiments in water where the flow is easy to visualize [43,44,48].

IV. HEAT TRANSPORT IN ROTATING CONVECTION

Heat transport measurements in the presence of rotation are complicated by the changing influences of buoyancy, proportional to Ra , and rotation, proportional to Ω_D . The simplest thing to do experimentally is to fix Ω_D and vary ΔT . Because the mean temperature changes with ΔT (fixed T_l), however, Ta changes owing to the temperature-dependent viscosity of water. This can be as large as 25% over the Ra range that we studied. As noted above σ also changes somewhat but that influence is small. Even constant Ta is not the appropriate variable to hold constant if one wants to evaluate the behavior of Nu as a function of Ra for a constant rotational forcing. From previous heat transport measurements [9], it appeared that Nu at fixed Ω_D was enhanced by rotation for intermediate Ra but seemed to asymptote to the nonrotating value of Nu at higher Ra . This situation was clarified by a numerical simulation of rotating convection [13,14] where a proper measure of rotational forcing is defined by the convective Rossby number Ro . $Ro \approx 1$ should mark the border between strongly rotating convection with $Ro \ll 1$ and weakly rotating convection with $Ro \gg 1$. In the numerical simulations of rotating convection at fixed $Ro = 0.75$, Nu scaled approximately as $Ra^{2/7}$. Here, we present heat transport measurements to higher Ra than previously [9] and consider Nu as a function of Ra at fixed Ro . In addition, we compare our results to earlier ones by Rossby [8] who compiled constant Nu contours as a function of Ra and Ta . To get a good understanding of the whole system, it is useful to consider different slices of the parameter space. We present them in the order of contours of Nu , fixed Ω_D , fixed Ra , and finally fixed Ro .

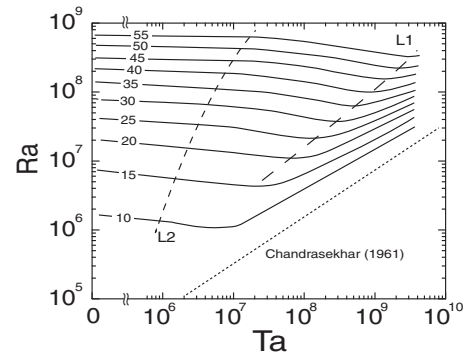


FIG. 4. Lines of constant Nu as a function of Ra and Ta . The dotted line is for the bulk convective onset from Chandrasekhar and $Ra_b \propto Ta^{2/3}$. Line $L1$ connects the loci of maximum Nu at constant Ra and agrees well with the loci of the minimum Ra at constant Nu . Line $L2$ delimits the parameter space into the right section where we made measurements under rotation and the left section, which are interpolated between zero rotation and the data with the lowest Ta .

Heat transport measurements in rotating convection can be summarized by a contour plot of Nu presented in Fig. 4. The points on each constant- Nu line were obtained by interpolating Nu data measured under different controlled conditions (namely, fixed Ω_D , Ra , and Ro). The individual data points are within 1% of the smooth curves for the largest $Nu \approx 55$ and within 3% for the smallest $Nu \approx 10$. Figure 4 complements the Nu contour plot in Fig. 11 of Rossby [8] where the highest Nu was 12, and the combination of the two gives a rather complete description of Nu in the parameter space of $Ra_b \leq Ra < 10^9$ and $0 \leq Ta < 10^{10}$. As shown in Fig. 11 of [8] for the lower Ra and Ta range, there is a minimum Ra for constant Nu or alternately there is a maximum Nu at constant Ra . At fixed Ta , however, Nu is a monotonically increasing function of Ra . In the following, results are presented which elucidate the origin of the maximum Nu at fixed Ra (or the minimum Ra at fixed Nu) and which investigate the variation in Nu as a function of Ra , Ta , and Ro . Before proceeding with these details, however, we can already see the overall trend of heat transport at fixed Ra . For low rotation, Nu is rather insensitive to changes in Ta (note the discontinuity in the horizontal axis in Fig. 4). In the intermediate range between lines $L2$ and $L1$, Nu increases with increasing Ta , as rotation enhances heat transport. For high enough Ta , however, rotation suppresses convection and Nu decreases as the onset of bulk convection is approached at $Ra_c(Ta)$.

A. Constant Ω_D

Shown in Fig. 5 is Nu versus Ra at $\Omega_D = 3.14$ rad/s where we obtained the background conductance K_b as described in Sec. II. The first few points have larger uncertainty because of the small temperature difference ΔT across the cell and the long thermal diffusion time in our cell (about 16 h). The transition to convection from a conduction state occurred at $\Delta T_c \approx 150$ mK and $Ra_c \approx 2 \times 10^6$, which is much lower than the theoretical value of 1.6×10^7 for a laterally infinite sys-

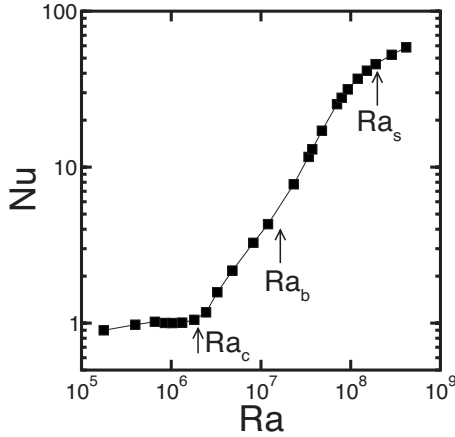


FIG. 5. Nu vs Ra for $\Omega_D=3.14$ rad/s ($Ta \approx 4.4 \times 10^9$). The arrows indicate the onset of the sidewall traveling state at Ra_c and the bulk state at Ra_b . The approximate onset of turbulent convection is indicated as Ra_s . The line is a guide to the eye.

tem at this rotation rate [7]. This lower-than-expected transition was observed in early heat transport experiments [8,9,31] and was later visually identified as a transition to a sidewall traveling-wave state [9,33]. Extrapolating the results in [9] to $Ta=3.3 \times 10^9$, one obtains the onset to the traveling state at about 3×10^6 , which is not far from our value of 2×10^6 . There is also an inflection point in Nu at $Ra_b \approx 1.6 \times 10^7$, coinciding with the theoretical prediction [7] for the transition to bulk convection [9,37]. Note that as the mean temperature of the cell increased during the measurements (constant T_r), Ta increased from 3.31×10^9 at small Ra to 5.12×10^9 at the highest Ra. We also measured Nu at fixed $\Omega_D=0.00, 0.188, 0.502$, and 1.26 rad/s. These measurements, shown in Fig. 6, served as a rough characterization of the system.

The first observation is that higher rotation suppresses convection relative to nonrotating convection from onset up to a value of Ra that depends on rotation. Above that Ra, Nu is higher than its corresponding value without rotation. Thus, the notion that rotation is a damping influence on convection,

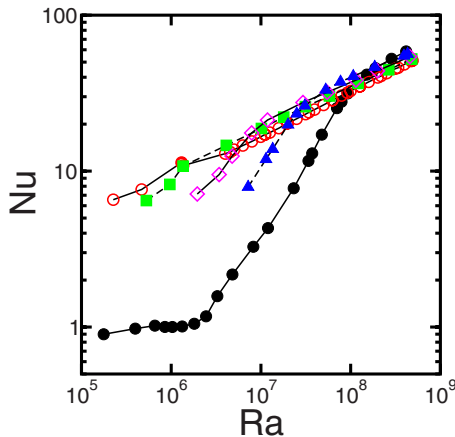


FIG. 6. (Color online) Nu vs Ra at constant Ω_D [approximate Ta]: 0.00 [0.0] (\circ), 0.188 [1.6×10^7] (\blacksquare), 0.502 [1.1×10^8] (\diamond), 1.26 [7.1×10^8] (\blacktriangle), and 3.14 [4.4×10^9] (\bullet). The lines are guides to the eye.

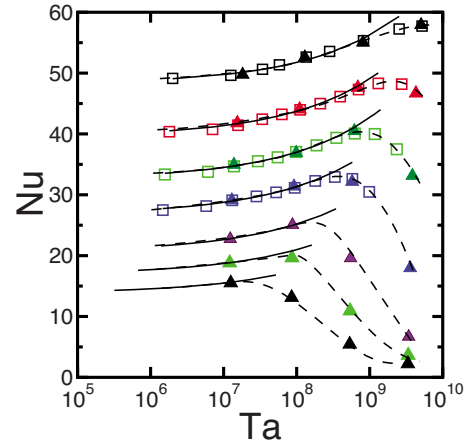


FIG. 7. (Color online) Nu vs Ta at (from top to bottom) $Ra = 4.0 \times 10^8, 2.0 \times 10^8, 1.0 \times 10^8, 5.0 \times 10^7, 2.0 \times 10^7, 1.0 \times 10^7$, and 5.0×10^6 . \square : measured experimentally with Ra held fixed and \blacktriangle : interpolated from data measured at constant Ω_D (Fig. 6). Solid lines are least-squares fits of the data to the form $Nu=Nu_0+\gamma Ta^{0.36}$; dashed lines are guides to the eye.

as suggested by the Taylor-Proudman theorem, is only valid near onset and the opposite is true for turbulent convection in this range of σ . The values of Ra_s for this crossover are plotted in Fig. 2 as solid circles. It is the crossover that gives rise to the maxima in the contours of Nu. The second thing to notice is that although $Nu(Ra, Ta)/Nu(Ra, 0) > 1$ at intermediate Ra, it appears to asymptote to 1 at higher Ra. This suggests that buoyancy wins out over rotation at high Ra and fixed Ta.

Revisiting the presence of mean flow, a feature of nonrotating convection over a large range of Ra, we consider mean flow for rotating convection. Flow visualization in the rotating frame for a cylindrical cell with $5 \times 10^7 < Ra < 5 \times 10^8$ [16] used both TLC and particle image velocimetry to determine the flow structure near the upper boundary layer [16]. The sheetlike plumes evolved under rotation into vortices and for small enough Ro, i.e., for rotation-dominated flow, there was no indication of a large-scale circulation extending over the size of the container. It seems that the shear on the boundary layer is of a very different form than for nonrotating convection as strong vortical motions dominate the flow just outside the boundary layer. Results for higher Ra in the range of 10^9 to 3×10^{11} [49] indicated a precessing mean flow provided $Ro > 0.5$. Similarly, recent results showed a breakdown of large-scale circulation for $Ro < 1.2$ [18].

B. Constant Ra

Another way to look at the influence of rotation on convection is to fix Ra and vary Ta. We measured Nu as a function of Ta at Rayleigh numbers of $Ra=5 \times 10^7, 1 \times 10^8, 2 \times 10^8, 4 \times 10^8$, with results as shown in Fig. 7. For each Ra, the data are corrected so that σ is constant. We have also included data for $Ra=2 \times 10^7, 1 \times 10^7, 5 \times 10^6$, which were obtained by interpolating data measured at constant Ro and Ω_D . For all the data sets, Nu increased with rotation before decreasing at higher Ta with a maximum Nu that var-

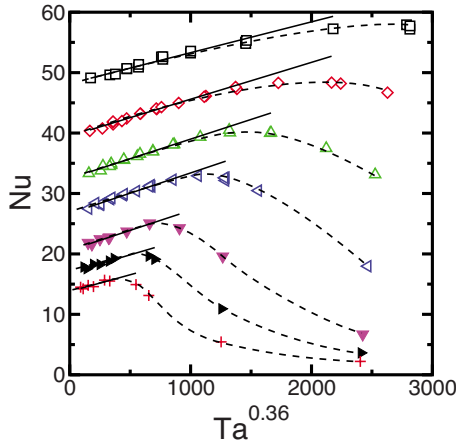


FIG. 8. (Color online) Nu vs $Ta^{0.36}$ so that least-squares fits (solid lines) of $Nu = Nu_0 + \gamma Ta^{0.36}$ yield straight lines. From top to bottom, $Ra = 4.0 \times 10^8$, 2.0×10^8 , 1.0×10^8 , 5.0×10^7 , 2.0×10^7 , 1.0×10^7 , and 5.0×10^6 . The dashed lines are guides to the eye and show deviation from linear fit.

ied with Ra . The loci Ta_m of maximum Nu are plotted in Fig. 2 as solid squares and in Fig. 4 (line $L1$). Note in the latter figure that this line approximately connects the loci of minimum Ra for each Nu contour. Line $L1$ can be approximated by $Ra \approx 2.2 Ta^{0.85}$, differing from the relation of $Ra \approx 206 Ta^{0.63}$ at lower Ra and Ta reported by Rossby (1969). The very different exponents indicate a continuously steepening curve and suggests that there is no clear asymptotic (in Ta) power-law scaling for the Nu maxima over the Ta range studied so far. Evaluating Nu along the line $L1$ yields the quantity Nu , discussed below.

An interesting conjecture regarding the enhancement of heat transport by rotation is that rotation creates thermal vortices which increase Nu through Ekman pumping in the boundary layer [9,14]. Thus, the enhancement in Nu might be proportional to the number of such vortices. Because we have not visualized the flow for the rotating system, the number of vortices as a function of Ta is not known directly [16,50]. Instead we consider the linear prediction for the number of structures at the convective onset. The linear wave number k_c increases with Ta and asymptotically scales like $Ta^{1/6}$ [7], which implies that the number of cellular structures should scale like $k_c^2 \sim Ta^{1/3}$. This scaling for vortex number was observed even significantly above onset in experiments with an open-top surface [34], which suggests that it is a reasonable assumption here. Instead of the $1/3$ scaling of the asymptotic theory, however, we will compare with an empirical fit to the linear data over our range of Ta , which gives $[k_c(Ta)/k_c(0)]^2 \approx 0.09 Ta^{0.36}$. In Fig. 8, we plot Nu in Fig. 7 as a function of $Ta^{0.36}$.

For $Ta < Ta_m$, there is a linear region for each Ra that shrinks as Ra decreases. In the linear region, we have

$$Nu = Nu_0 + \gamma Ta^{0.36}, \quad (4)$$

where Nu_0 and γ are fitting parameters and values for different Ra are listed in Table II. Solid lines in Figs. 7 and 8 are calculated from these fitting parameters. The deviation from

TABLE II. Fitting parameters Nu_0 and γ . Nu_m and Ta_m are the maximum Nusselt number and its location at constant Ra . $Nu(0)$ is the nonrotating value obtained by cubic least-squares interpolation of $\log_{10}(Nu)$ vs $\log_{10}(Ra)$ for nonrotating convection.

Ra ($\times 10^6$)	Nu_m	Ta_m ($\times 10^7$)	$Nu(0)$	Nu_0	γ
400	58.0	420	48.38	48.2(2)	0.0051(5)
200	49.0	200	39.87	39.5(1)	0.0061(4)
100	40.5	90	32.62	32.5(1)	0.0065(2)
50	33.0	42	26.60	26.7(1)	0.0067(3)
20	25.5	14	20.31	20.6(2)	0.0066(4)
10	20.5	6	16.62	16.9(2)	0.0059(9)
5	16.0	2.5	13.65	13.8(3)	0.0055(12)

the solid lines (linear behavior) at higher Ta is a result of rotation suppressing convection in the weakly nonlinear regime near onset. Also listed in Table II is $Nu(\Omega_D=0)$, the Nusselt number of nonrotating convection. The fitting parameter Nu_0 for all Ra is nearly identical to $Nu(0)$ within fitting and experimental uncertainties. This indicates that the fitting is consistent with the data in the range $0 < Ta < Ta_m$. Thus, the enhancement of Nu by rotation is given by $\Delta Nu = \gamma Ta^{0.36}$ with γ about 0.006. This result supports the conjecture that the enhancement is proportional to the average number of thermal vortices.

Using velocity field measurements [16], the number of vortices could be evaluated more quantitatively. At a fixed $Ra = 3.2 \times 10^8$, Ta was varied from 0 up to about 10^{10} . The variation in vortex (cyclonic) density with $\Omega \sim Ta^{1/2}$ was slightly sublinear, over the range $10^7 < Ta < 10^9$. This result suggests that the number of vortices depends on Ta with a power a bit less than $1/2$, which is roughly consistent with the estimate based on linear stability arguments. Unfortunately, the vortex density data are too sparse to provide better estimates in the Ta range of interest.

C. Constant Ro : Power-law scaling

Many experiments in thermal convection without rotation showed that Nu scales more closely as the $2/7$ power of Ra in the regime $10^7 < Ra < 10^9$ than with the classical $1/3$ power law. As discussed earlier, a generalized theory in terms of a phase diagram in Ra and σ and more precise experimental measurements suggest a form with the sum of two power laws with exponents of $1/5$, $1/4$, $1/3$, or $1/2$ depending on the region in phase space. Rotation complicates the issue of scaling since the relative influence of rotation changes with changing Ra at fixed Ta . Numerical simulations [14] showed that the convective Rossby number is a good measure of the relative importance of buoyancy with respect to rotation: for $Ro = 0.75$ and $\sigma = 1$, they found that $Nu \sim Ra^{2/7}$, which indicates that the details of rotation are relatively unimportant in the determination of the scaling exponent. We have tested this result and over the range $0.1 < Ro < 1.5$, we also find approximate $2/7$ power-law scaling. We use this single power-law description for convenience—a fit of the form

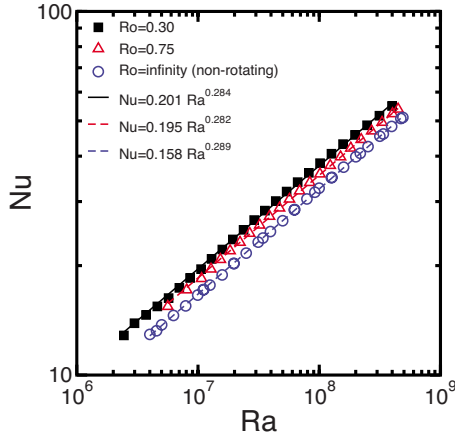


FIG. 9. (Color online) Nu vs Ra at constant Ro: 0.30 (■), 0.75 (△), and ∞ (○). The dashed lines are power-law fits with amplitudes and exponents listed in the legend.

$Nu = a Ra^{1/5} + b Ra^{1/3}$ yields equivalent fits. For the nonrotating case, the coefficients are $a=0.26$ and $b=0.04$, for $Ro=0.75$, $a=0.33$, and $b=0.048$ and, for $Ro=0.30$, $a=0.33$, and $b=0.053$.

We have measured Nu at several Rossby numbers: $Ro = 0.12, 0.30, 0.52, 0.75, 1.15, 1.49$, and ∞ (zero rotation). For clarity, only parts of the data for $Ro=0.30, 0.75$, and ∞ are plotted on a log-log scale in Fig. 9 with the coefficients of power-law fits listed in the plot. The Nusselt number agrees well with a 2/7 power law and is different from the 1/3 power law, especially at higher Rayleigh number ($>4 \times 10^7$). The coefficient $A_{2/7}$ is plotted in Fig. 10(a). The trend of decreasing $A_{2/7}$ with increasing Ro demonstrates the enhancement of heat transport by rotation. An empirical fit to the coefficient $A_{2/7}$ yields its approximate Ro dependence: $A_{2/7} = 0.17(1 + 0.24e^{-Ro/0.8})$.

We also fit the data with $Nu = A_\beta Ra^\beta$ to obtain the coefficient and the exponent as functions of Ro. Least-squares fitting was performed in the range $4 \times 10^7 < Ra < 5 \times 10^8$ to avoid possible deviation from power-law scaling at lower Ra.

In Fig. 11, we plot Nu/Ra^β versus Ra to gauge how well the power law with exponent β describes the data. Over the Ra range where the a power law with exponent β is satisfied, a constant value of $A_\beta = Nu/Ra^\beta$ is expected. The description is reasonably good for the nonrotating case, but becomes a little worse as Ro decreases or rotation increases. In all three cases, the value of A_β varies by less than $\pm 2\%$ over the range of Ra. The results for A_β and β versus Ro are plotted in Figs. 10(b) and 10(c), respectively. For nonrotating convection the coefficient $A=0.164$ and exponent $\beta=0.286$ are slightly different from the values obtained by fitting the data for $4 \times 10^6 < Ra < 5 \times 10^8$. For finite Ro, the exponent depends on Ro: it increases almost monotonically from 0.269 at $Ro=0.12$ to 0.287 at $Ro=1.5$. As shown in Figs. 9–11, the coefficient A decreases as Ro increases (lower rotation) although its value is sensitive to the fitting value of the exponent as can be seen by comparing Figs. 10(a) and 10(b).

Determining unambiguously the scaling behavior of the heat transport requires many orders of magnitude in Ra. Thus, an absolute comparison of scaling exponents in our

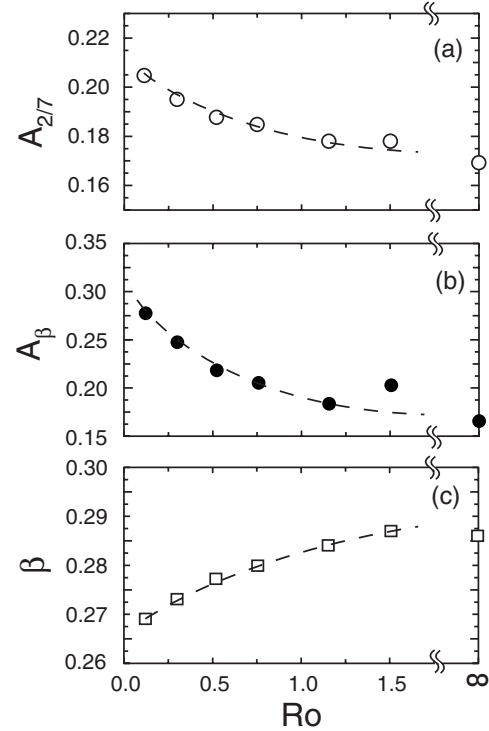


FIG. 10. Rotation-dependence of the fitting parameters: (a) coefficient $A_{2/7}$ in $Nu = A_{2/7} Ra^{2/7}$, (b) A_β , and (c) β in $Nu = A_\beta Ra^\beta$. Dashed line in (a) is fit $A_{2/7} = 0.17(1 + 0.24e^{-Ro/0.8})$, whereas others in (b) and (c) are guides to the eye.

experiment is uncertain. Nonetheless, our experiments yielded some interesting results, especially when compared to numerical simulations [13,14]. First, at fixed Ro, Nu depends on Ra with a power law close to 2/7, in agreement with numerical simulation [14] for $Ro=0.75$ and $\sigma=1$. If anything, rotation seems to reduce the scaling exponent slightly. This could be the result of different scaling ranges as a function of Ro, because fixed Ro does not exactly maintain a balance between buoyancy and rotation, or because rotation modifies the scaling exponent directly. An extended range in Ra would be necessary to resolve this quantitatively, perhaps in a gas system.

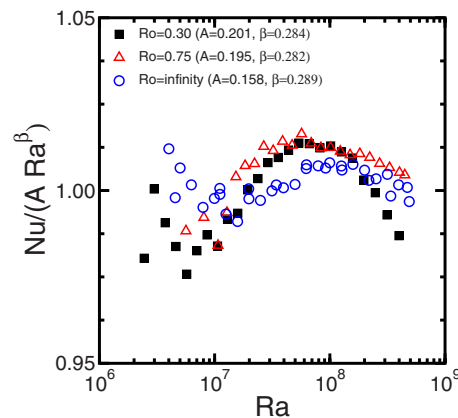


FIG. 11. (Color online) Nu/Ra^β vs Ra at constant Ro: 0.30 (■), 0.75 (△), and (○, nonrotating).

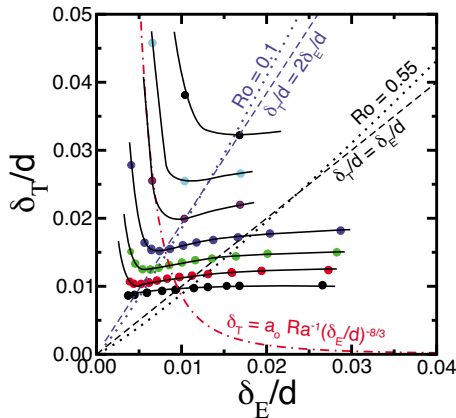


FIG. 12. (Color online) Calculated thermal boundary layer thickness δ_T vs Ekman layer thickness δ_E at (from top to bottom) $Ra=4.0 \times 10^8$, 2.0×10^8 , 1.0×10^8 , 5.0×10^7 , 2.0×10^7 , 1.0×10^7 , and 5.0×10^6 . Same data as in Fig. 7. Solid lines are guides to the eye. The bottom [top] dashed line labeled in the plot shows the condition $\delta_T = \delta_E$ [$\delta_T = 2\delta_E$]. The bottom (top) dotted line labeled in the plot shows the lines of constant $Ro=0.55$ (0.1). The dashed-dotted line shows a contour where $Nu \sim Ra$, a condition which translates to $\delta_T/d = (a_o/Ra)(\delta_E/d)^{-8/3}$ with $a_o \sim 1$.

From the perspective of turbulent convection theory [26,28,51], the relative insensitivity of the scaling exponent to rotation is rather interesting because rotation affects many properties of the turbulence, such as the change from thermal plumes to vortices and the existence of a turbulent Ekman boundary layer and associated Ekman pumping. In nonrotating convection, the relationship between the thermal boundary layer thickness δ_T and the viscous sublayer thickness δ_ν determines the power-law scaling in the sheared boundary layer theory [28] where $Nu \sim Ra^{2/7}$ applies when $\delta_T < \delta_\nu$. Rotation introduces another vertical length scale, the Ekman layer thickness δ_E , which could, in principle, play a role similar to δ_ν in nonrotating convection.

Numerical simulations [13] showed that the transition to turbulent scaling occurs at $Ra \approx 4 \times 10^7$ for $Ro=0.75$, where $\delta_T \approx \delta_E$. Further suggestion that this is a more general condition was presented in an analysis of experimental and numerical data [19]. We can use our data to indirectly consider the relationship between thermal and Ekman boundary layer thicknesses for different conditions of the flow. A direct measurement of local thermal and velocity boundary layers has not been done here. Rather, we consider the quantities δ_T and δ_E defined as

$$\delta_T = \frac{1}{2} \frac{d}{Nu}, \quad (5)$$

$$\delta_E = \left(\frac{\nu}{2\Omega_D} \right)^{1/2} = \frac{d}{Ta^{1/4}} = d Ro^{1/2} \sigma^{1/4} Ra^{-1/4}. \quad (6)$$

In Fig. 12, we plot δ_T versus δ_E at different Ra using the same data as in Fig. 7. The data have the general form that at high δ_E/d the flow is turbulent and δ_T/d decreases slightly with decreasing δ_E/d indicating the trend of slightly increased Nu (smaller δ_T/d) at higher Ta (smaller δ_E/d). From

the analysis above for Fig. 7, we have $Nu = Nu_0(Ra) + \gamma Ta^{0.36}$, which implies the relationship $\delta_T/d = \delta_{T_o}/d / (1 + \gamma Ta^{0.36} d / 2\delta_{T_o}) = \delta_{T_o}/d / [1 + \gamma (\delta_E/d)^{-1.4} d / 2\delta_{T_o}]$, where $\delta_{T_o} = 1/[2Nu_0(Ra)]$. In this regime Nu scales approximately as $Ra^{2/7}$ with a Ta -dependent prefactor as discussed above. As δ_E/d decreases further, one approaches a transition to a regime with approximately linear scaling such that $Nu \sim Ra/Ra_b$ where we have that $Ra_b \sim Ta^{2/3}$. (In [19], a slightly different form $Nu \sim (Ra/Ra_b)^{6/5}$ was used.) Translating this expression, we have $\delta_T/d = (a_o/Ra)(\delta_E/d)^{-8/3}$ (bounded above by the condition $\delta_T/d = 0.5$), where a_o is an order one coefficient. This curve is plotted in Fig. 12 and reproduces well the rapid increase in δ_T/d for small δ_E/d .

There are several other interesting comparisons to make for these data. The transition from linear scaling to sublinear turbulent scaling occurs along a line given by $\delta_T/d \approx 2\delta_E/d$ shown in Fig. 12. This line is quite close to the condition $Ro=0.1$, also shown as a dotted line. On the other hand the condition $\delta_T/d = \delta_E/d$ corresponds to constant $Ro=0.55$ as illustrated by curves in the plot. Thus, the condition postulated recently [19] that the transition to turbulence occurs when $\delta_T/d \approx \delta_E/d$ as suggested earlier from numerical simulations [13] is not quantitatively confirmed for our data. Rather, the condition $Ro \approx 0.1$, or equivalently $\delta_T/d \approx 2\delta_E/d$, seems to more accurately describe the transition to the turbulent regime in rotating convection than does the condition $\delta_T/d \approx \delta_E/d$. Whether this factor of two is an important distinction remains unclear.

In summary, the interplay of these two length scales is not yet understood based on measurements done up to this point; it is not really even clear that δ_E defined from nonconvecting problems with differential rotation is the proper variable to use here. A direct measurement of the turbulent Ekman layer in the presence of a thermal boundary layer would be very useful to augment the arguments based on numerical simulations [13] for a coexisting thermal boundary layer with a linear Ekman layer. Subsequent flow visualization [16] using particle image velocimetry showed a diverse set of interesting behaviors of velocity and vorticity fluctuations but did not yield a definitive conclusion regarding the complex interplay of thermal and kinetic boundary layers involved in determining heat transport for turbulent rotating convection. Recent measurements of kinetic boundary layers [18] may help to resolve these questions as will pushing to higher Ra to see if the trends presented here continue.

V. CONCLUSIONS

We have presented experimental studies of turbulent thermal convection in water confined in a cell with a square cross section with and without rotation [36]. In nonrotating convection, the Nusselt number was found to scale roughly as $Ra^{2/7}$ for $4 \times 10^6 < Ra < 5 \times 10^8$. Heat transport measurement in rotating convection confirmed the findings by other researchers [8,9] that rotation enhances thermal transport over a certain range of Ra and Ta for $\sigma \approx 6$. This enhancement was attributed to Ekman pumping of the boundary layer by rotation-induced vortices [13,52] and was supported in this study by results that showed indirectly that the en-

hanced heat transport was roughly proportional to the number of vortex structures—each contributing an extra bit of Ekman pumping of the boundary layer. At fixed Rossby number, Nu was found to scale approximately as $Ra^{2/7}$, as does Nu of nonrotating convection and predicted from numerical simulation [14]. Characterization by a combination of power laws [26,27] was equally good at fitting the data for both rotating and nonrotating convections.

Analysis using Ekman layers instead of kinetic boundary layers as input into a scaling theory did not provide additional insight into the heat transport data, and it remains unclear how rotation and its associated modification of boundary layer structure affects heat transport. Given the extensive study of nonrotating convection in recent years with great advances in characterizing boundary layers, heat transport, and large-scale circulation [5], there seems to be an emerging opportunity to apply similar rigor to the geophysically important case of rotating thermal convection [18,17]. We hope

that our work may stimulate more experimental and theoretical studies on rotating thermal turbulence. A rigorous test of the power-law scaling of Nu under constant Ro and the validity of Ro as the “good” parameter in rotating convection requires a much larger Ra and Ta range than what was available in our experiment. A gas convection may be needed to achieve such experimental conditions. The interplay among various length scales, such as horizontal vortex wavelength, Ekman layer, thermal boundary layer, and viscous sublayer, was not well understood in our experiment and certainly calls for further work on this aspect of turbulent convection.

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